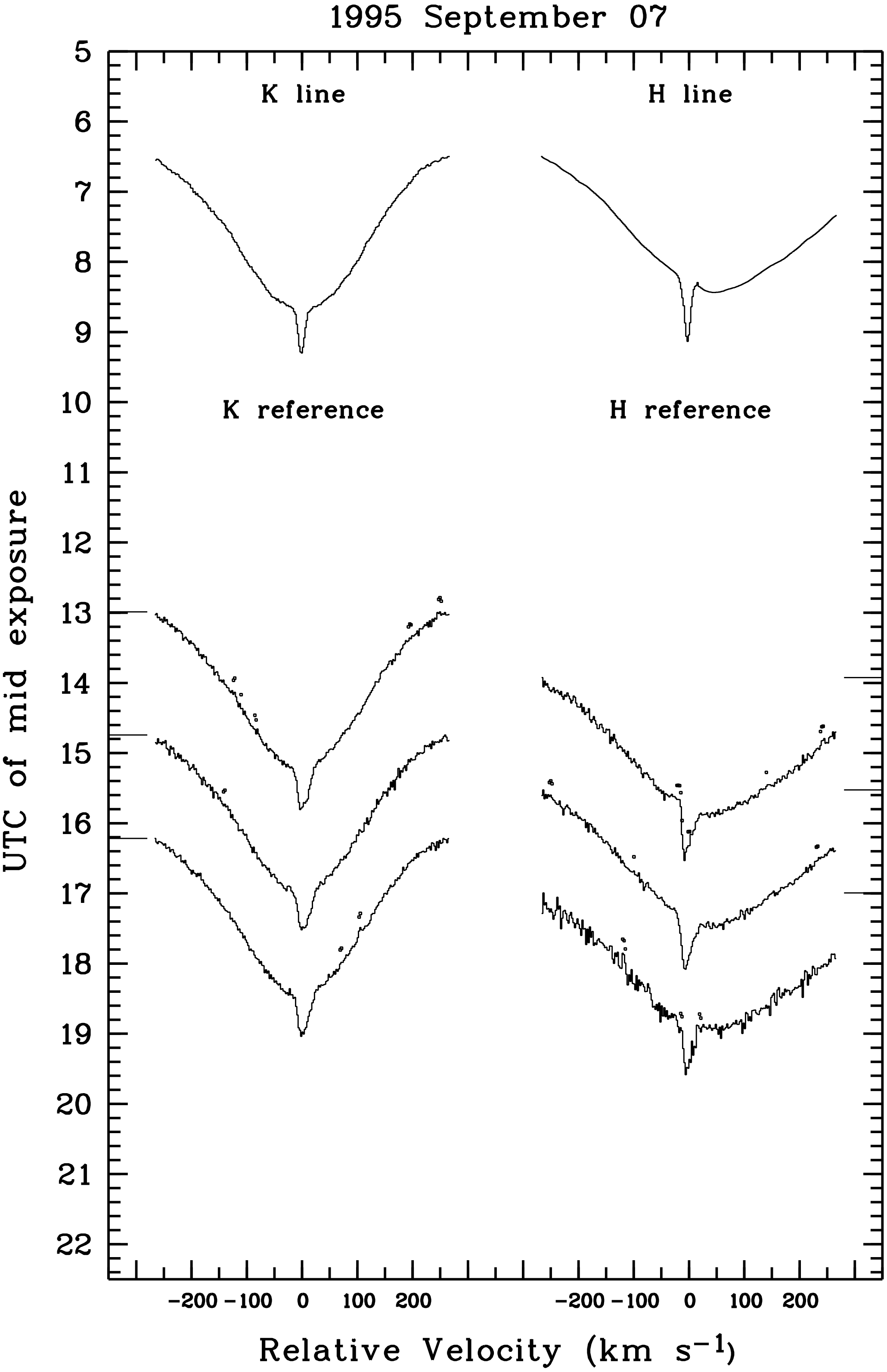

Figure 22: Observations of β Pictoris on 1995 September 7 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

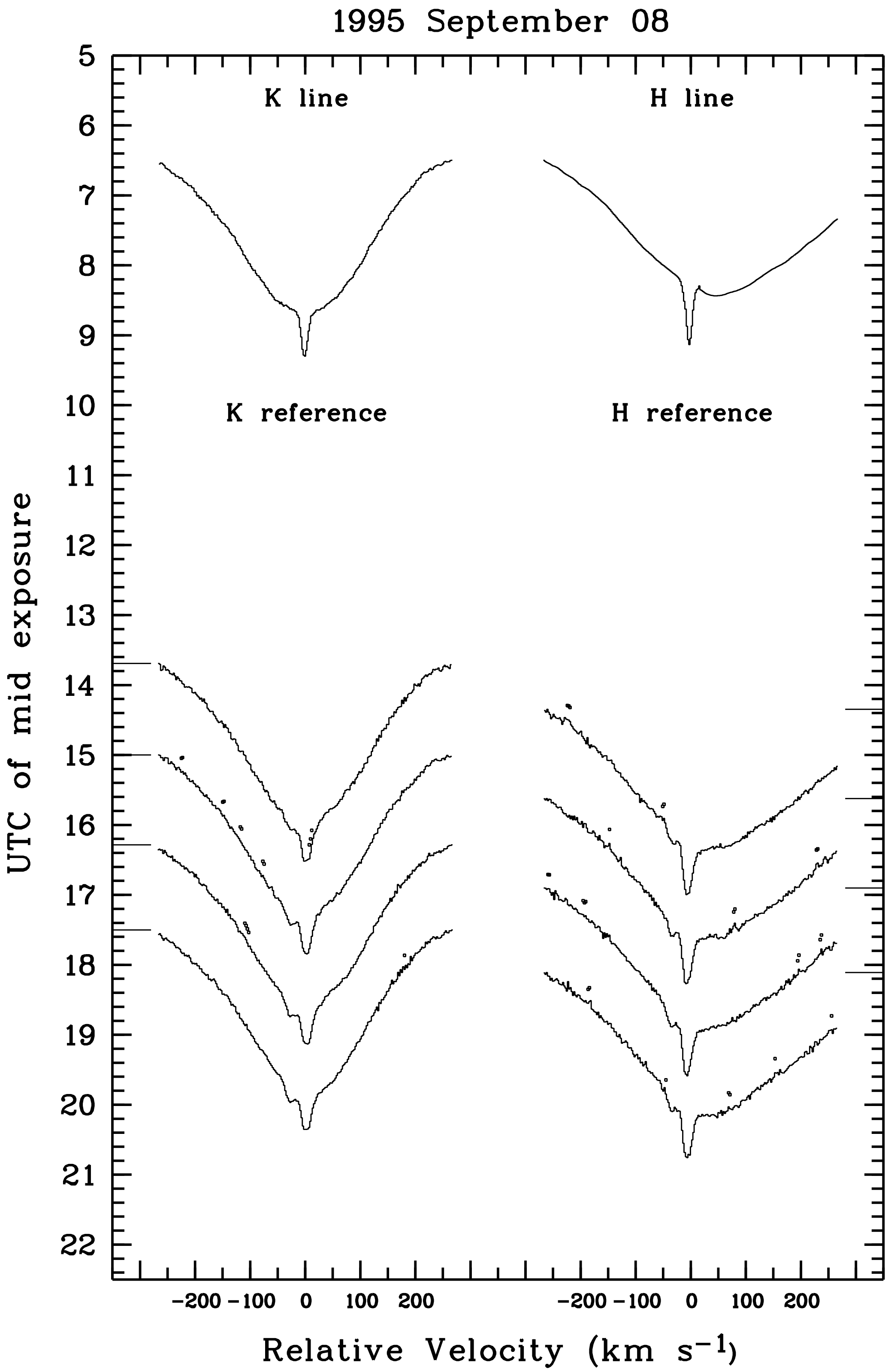

Figure 23: Observations of β Pictoris on 1995 September 8 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

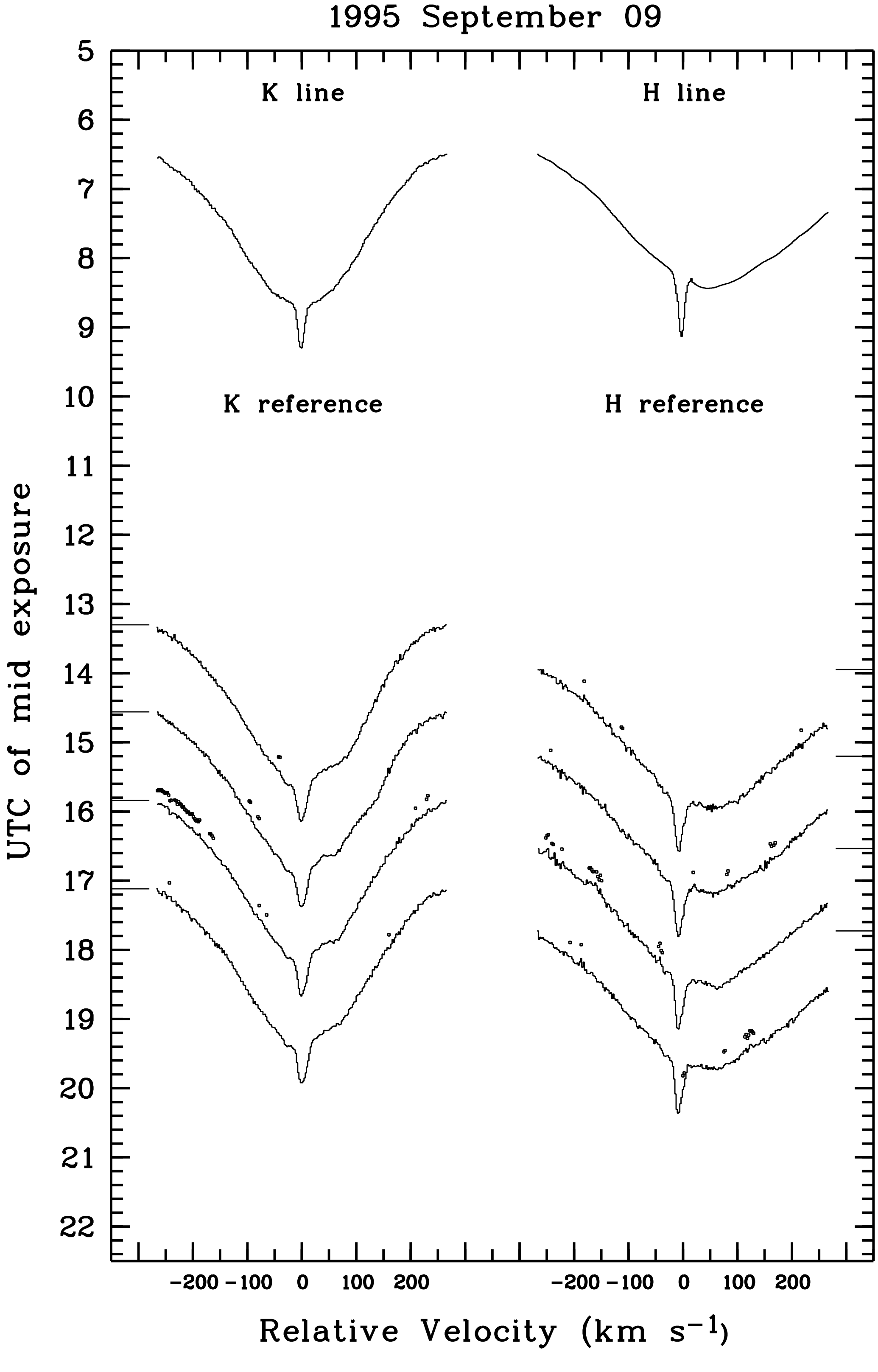

Figure 24: Observations of β Pictoris on 1995 September 9 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

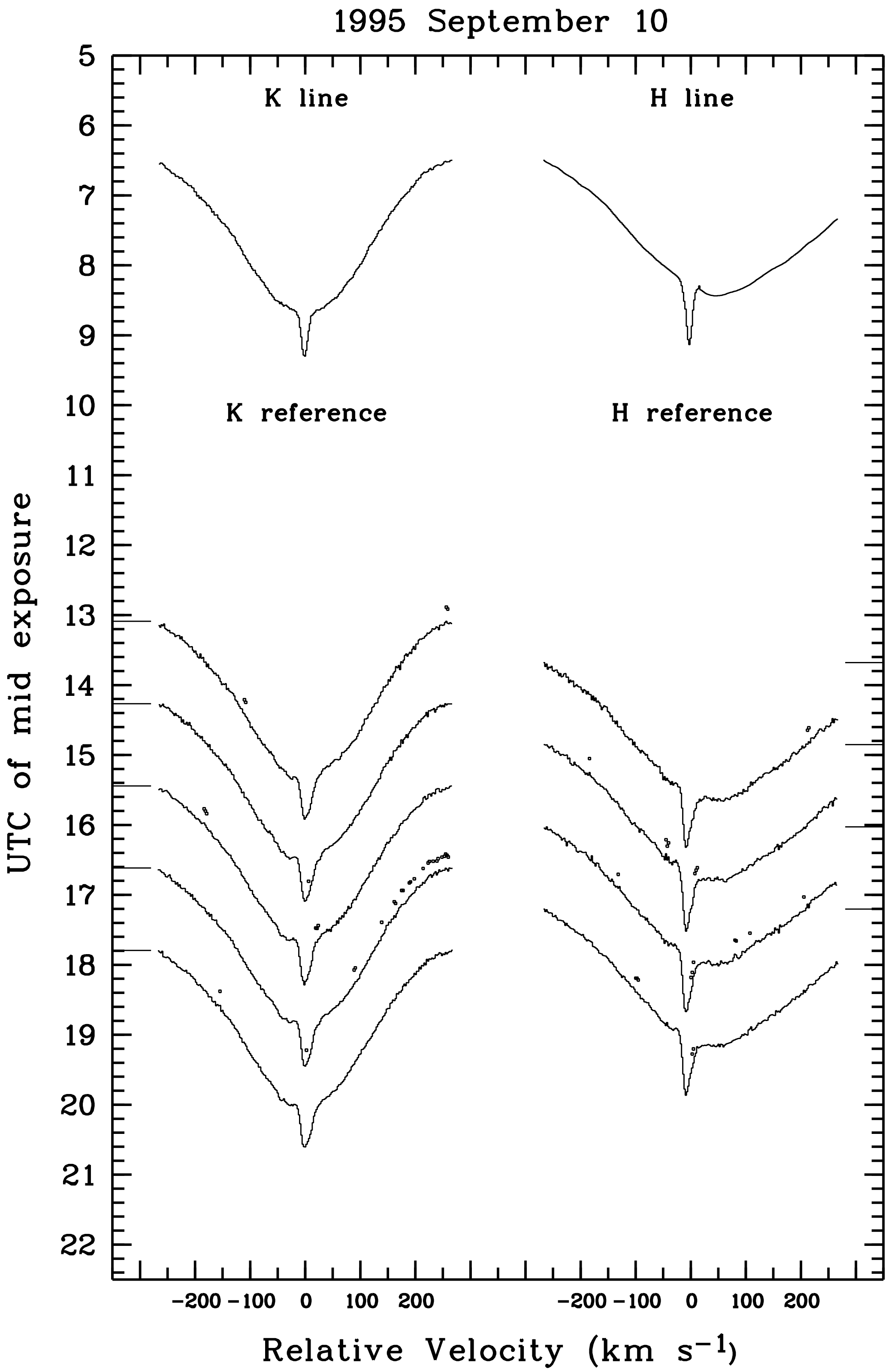

Figure 25: Observations of β Pictoris on 1995 September 10 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

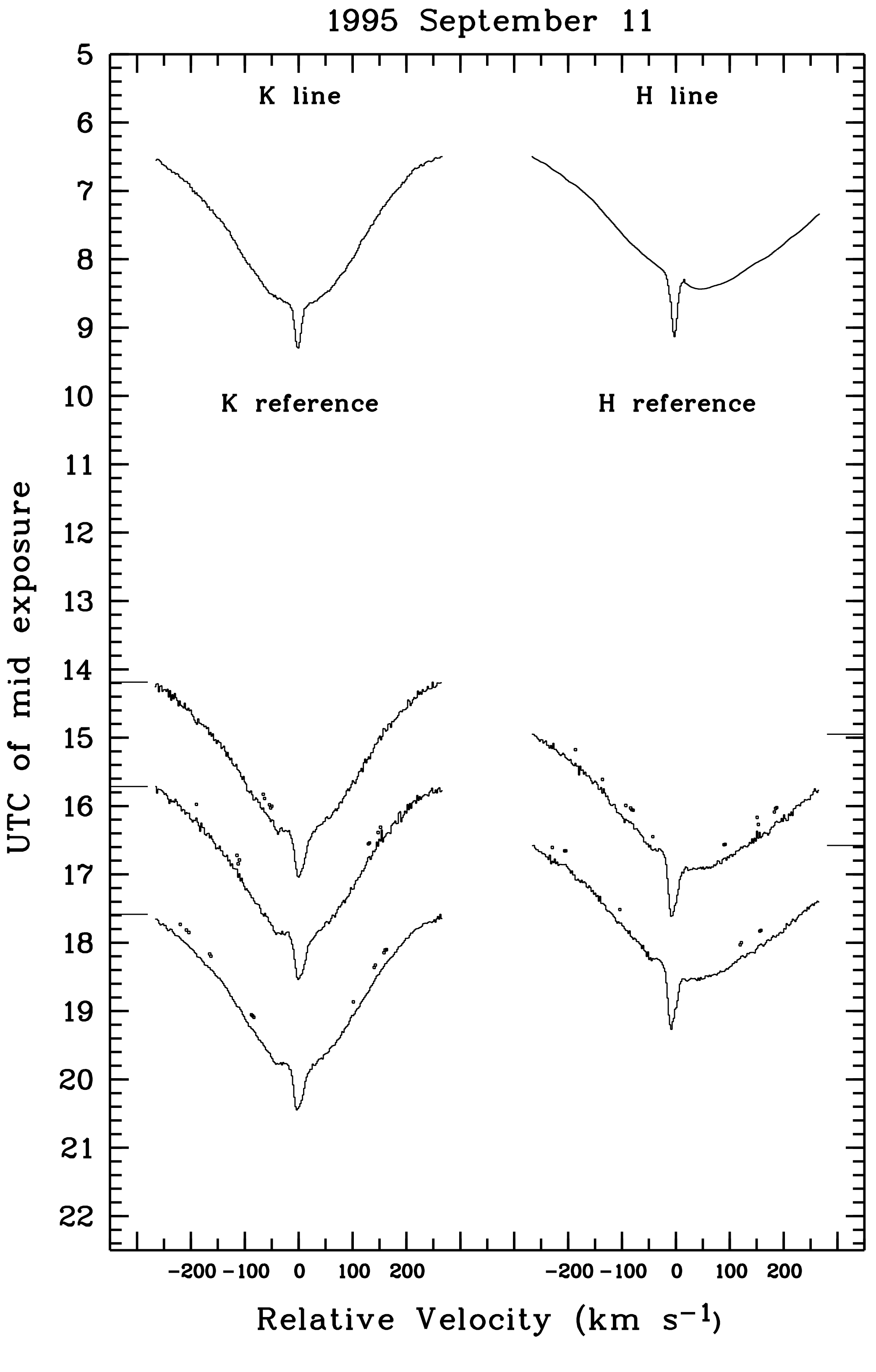

Figure 26: Observations of β Pictoris on 1995 September 11 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

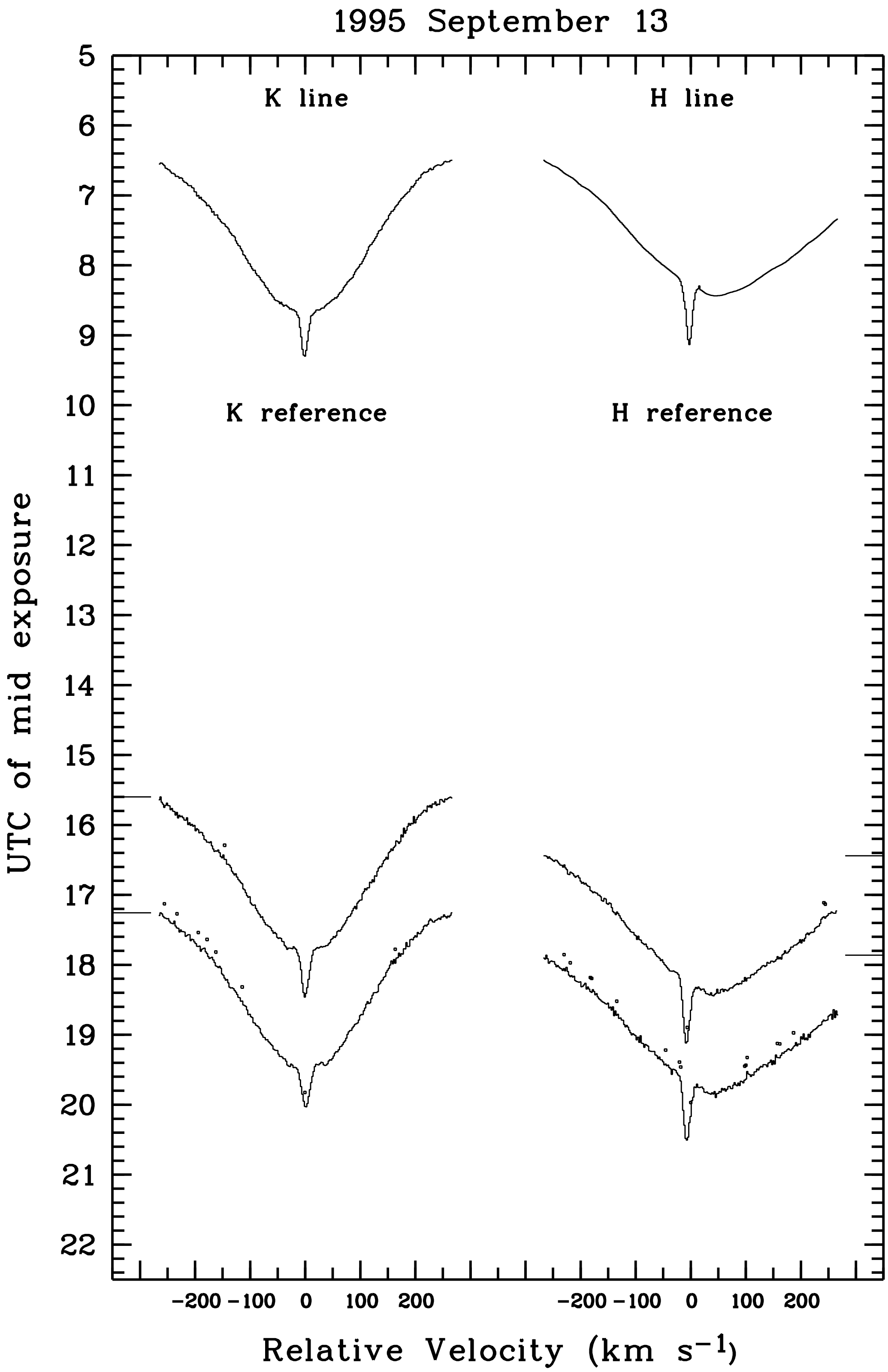

Figure 27: Observations of β Pictoris on 1995 September 13 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

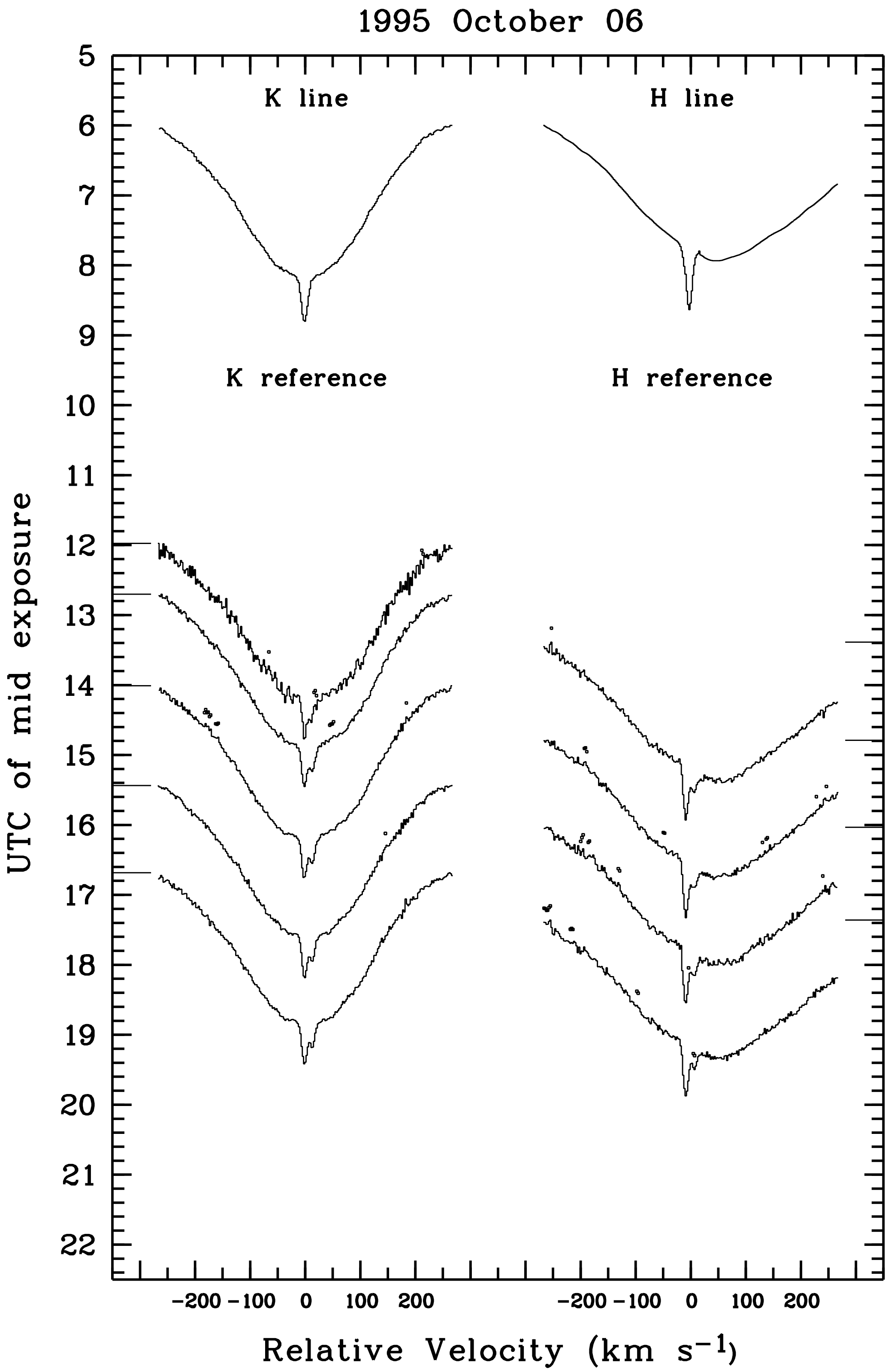

Figure 28: Observations of β Pictoris on 1995 October 6 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

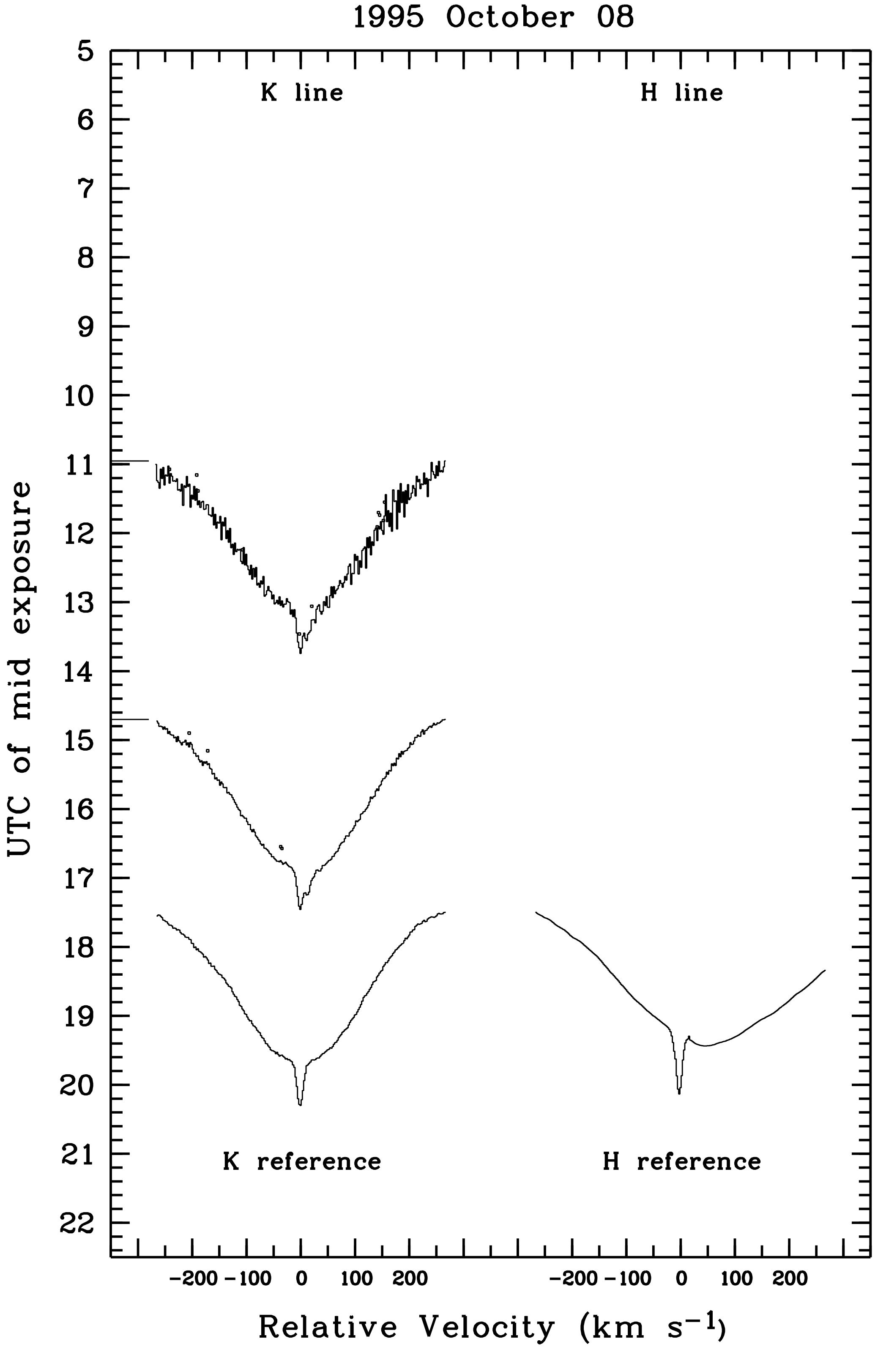

Figure 29: Observations of β Pictoris on 1995 October 8 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

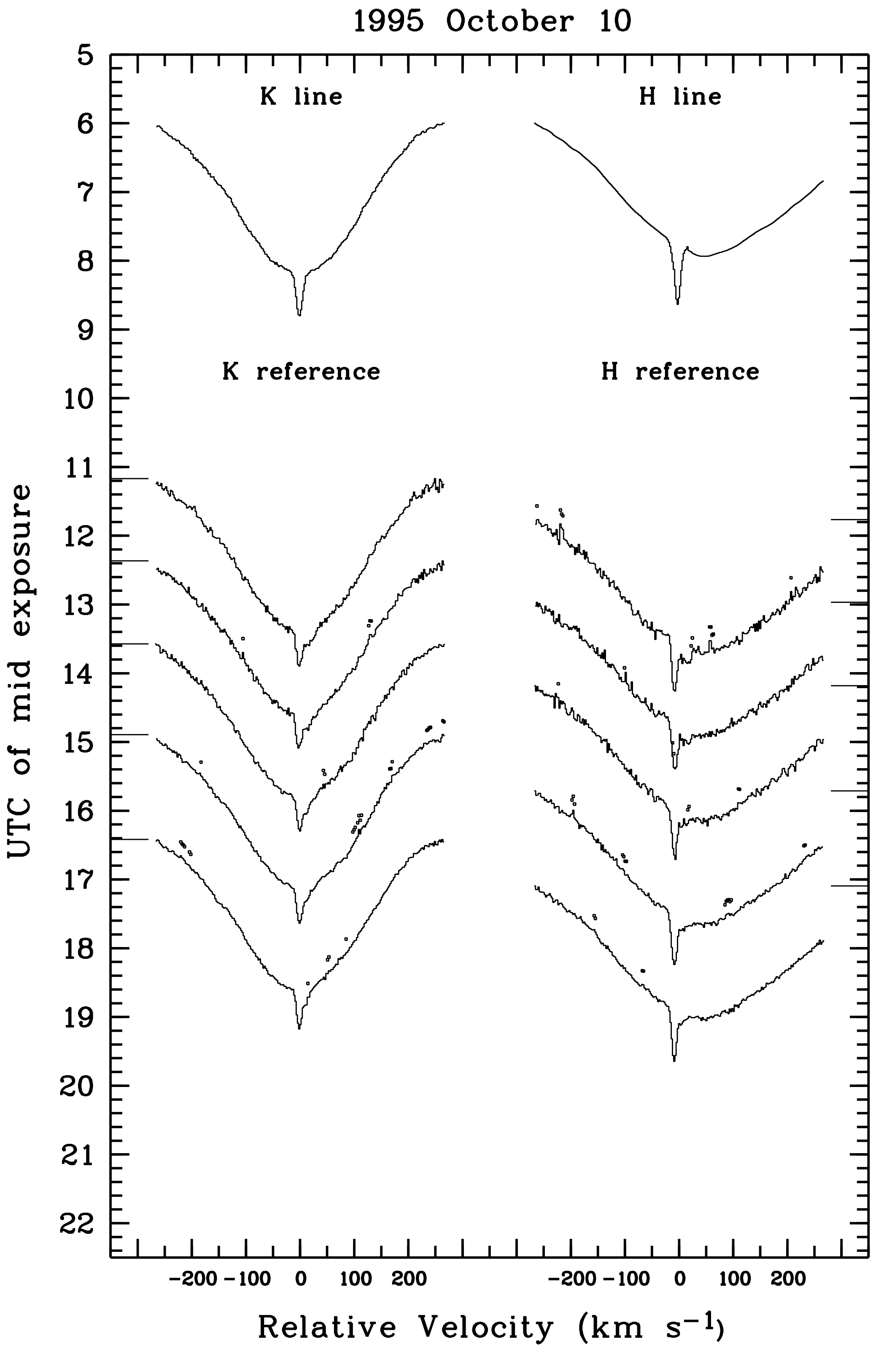

Figure 30: Observations of β Pictoris on 1995 October 10 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

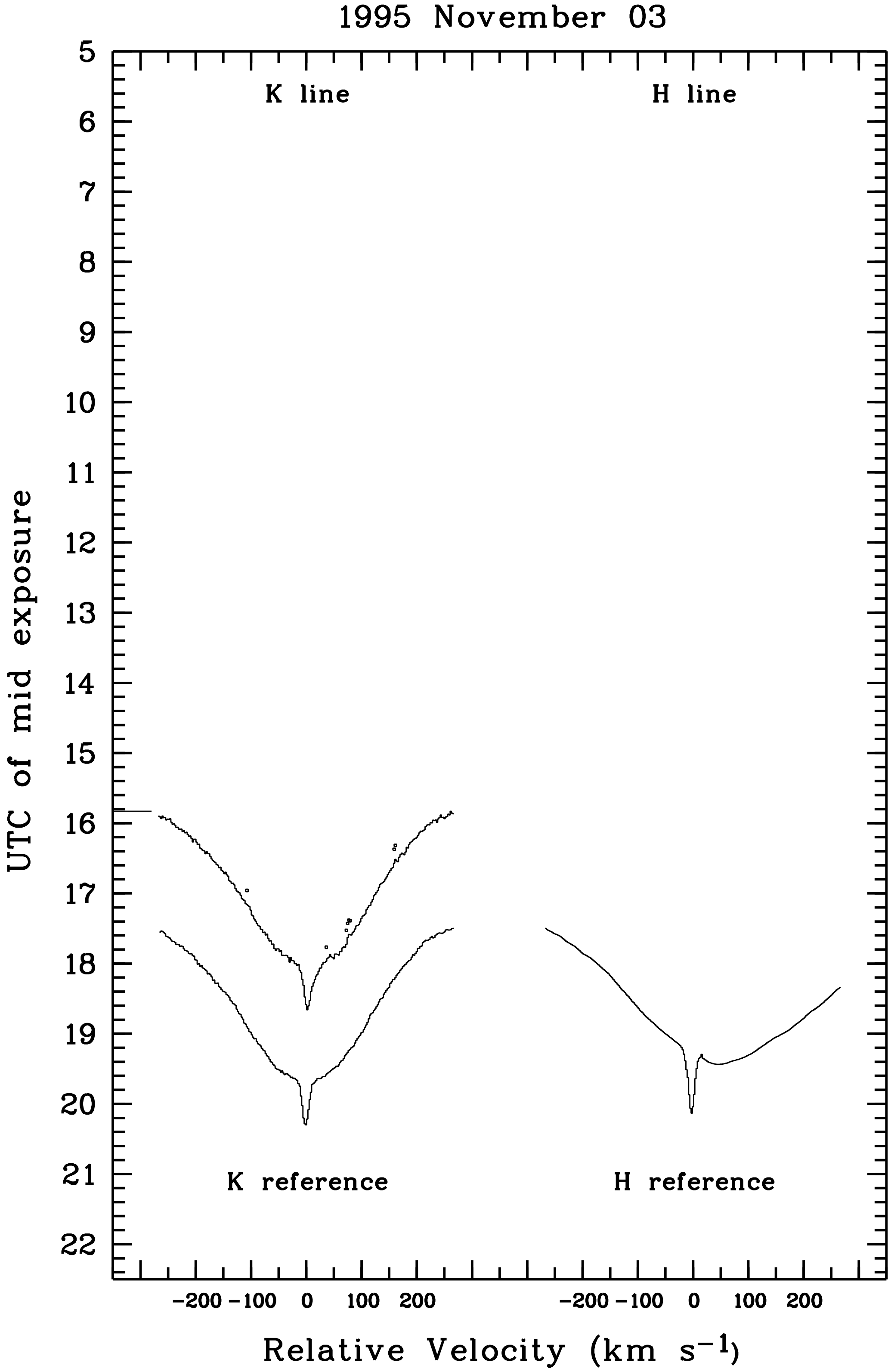

Figure 31: Observations of β Pictoris on 1995 November 3 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

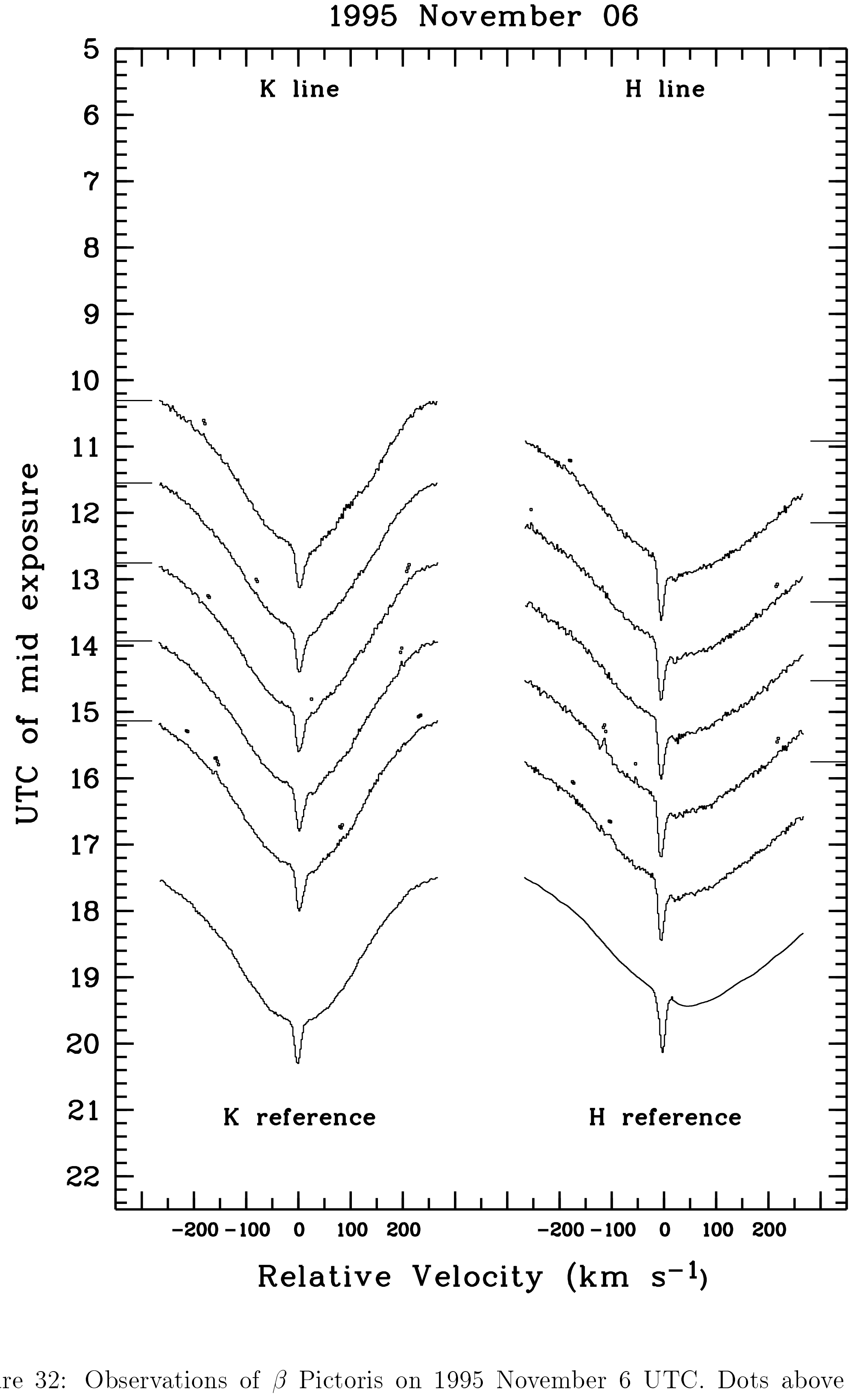

Figure 32: Observations of β Pictoris on 1995 November 6 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

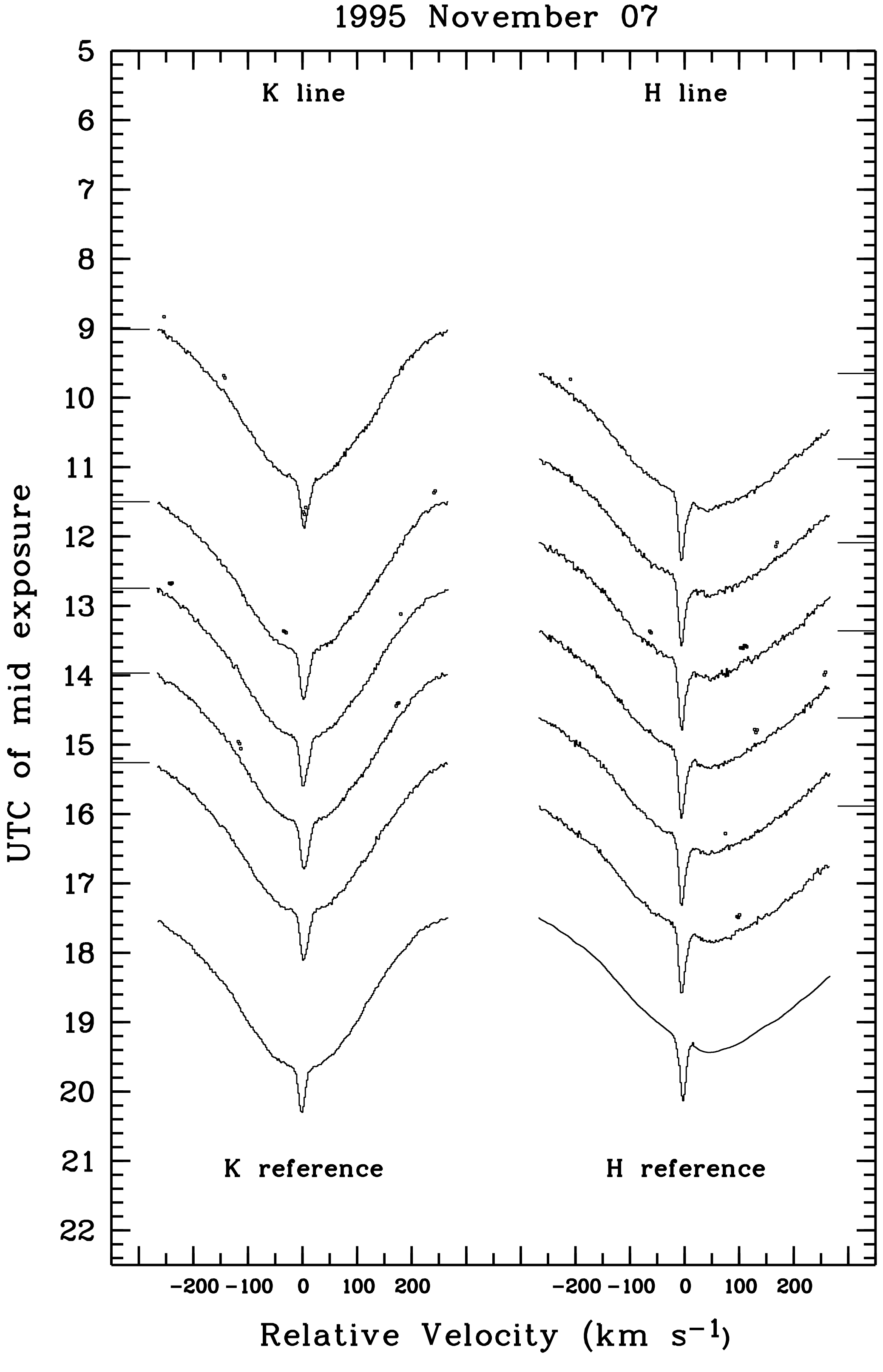

Figure 33: Observations of β Pictoris on 1995 November 7 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

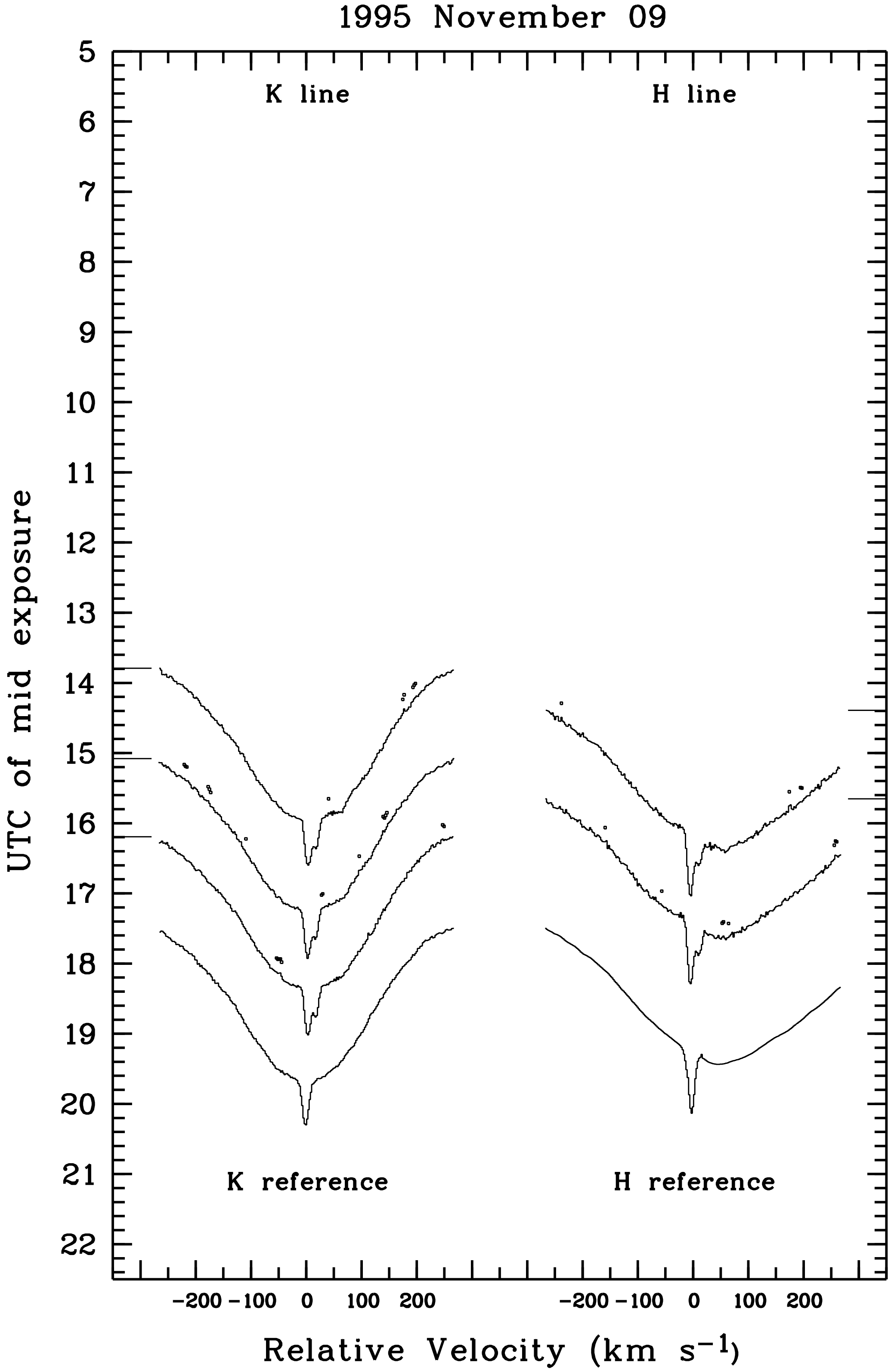

Figure 34: Observations of β Pictoris on 1995 November 9 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

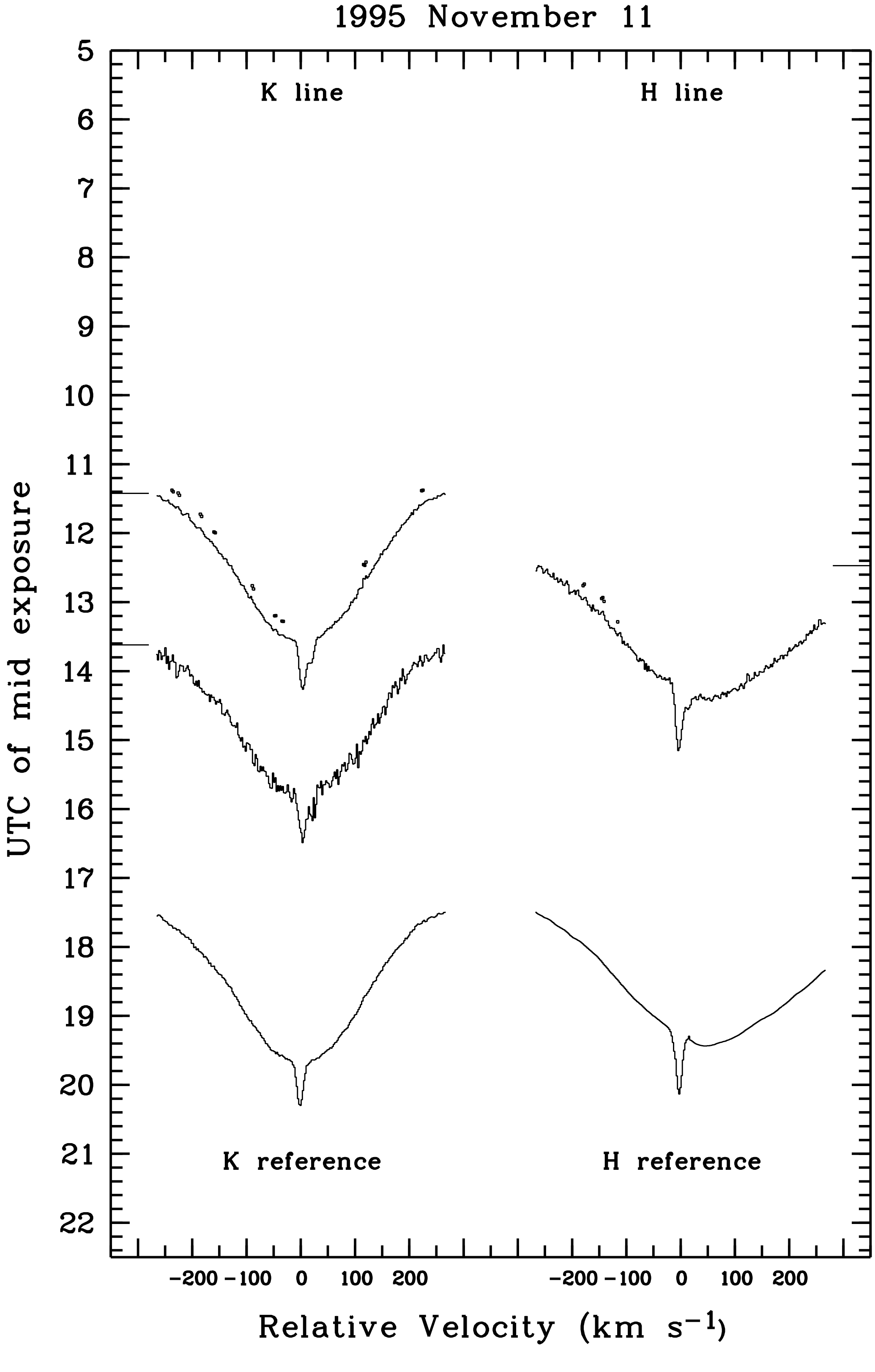

Figure 35: Observations of β Pictoris on 1995 November 11 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

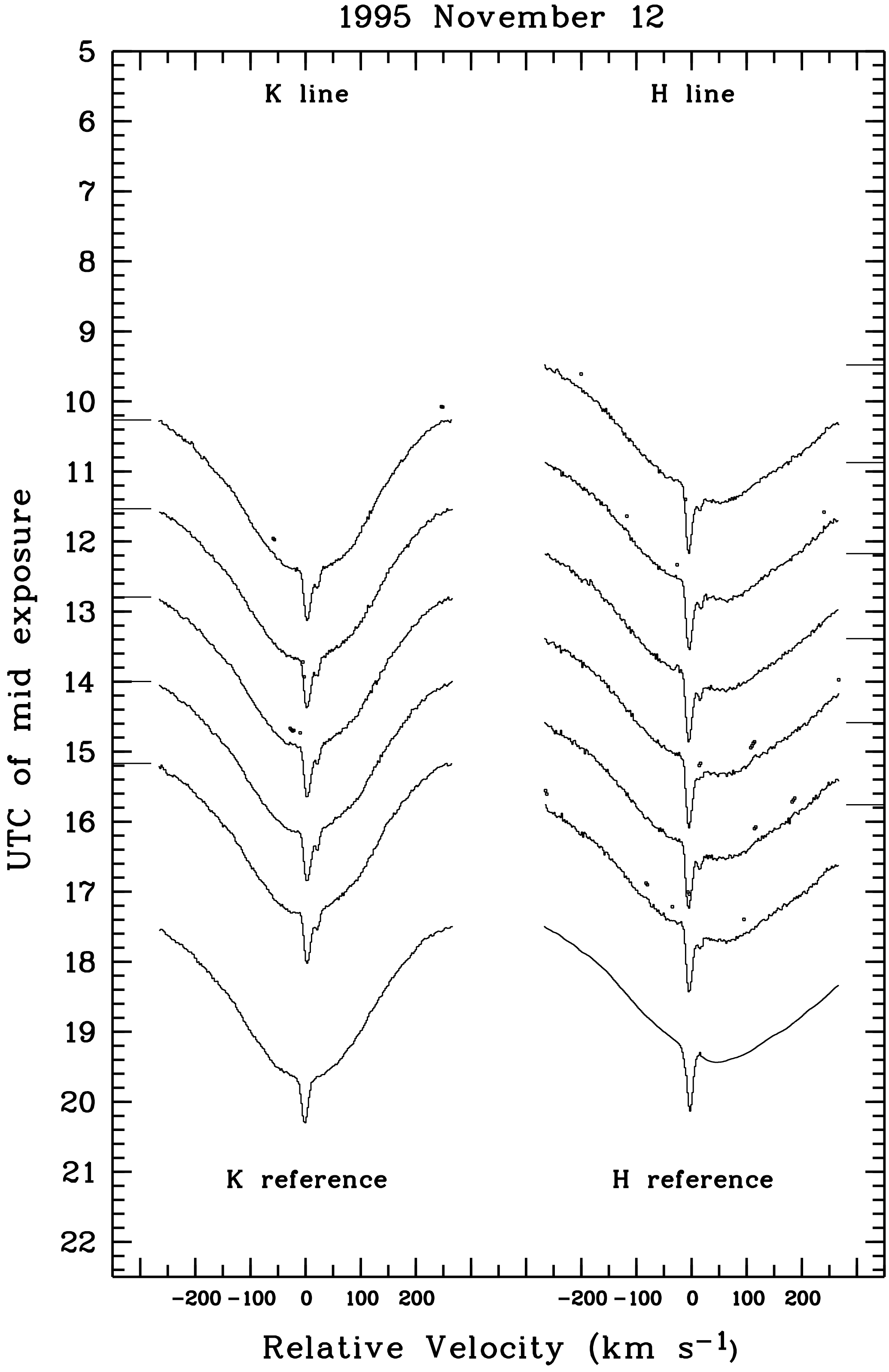

Figure 36: Observations of β Pictoris on 1995 November 12 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

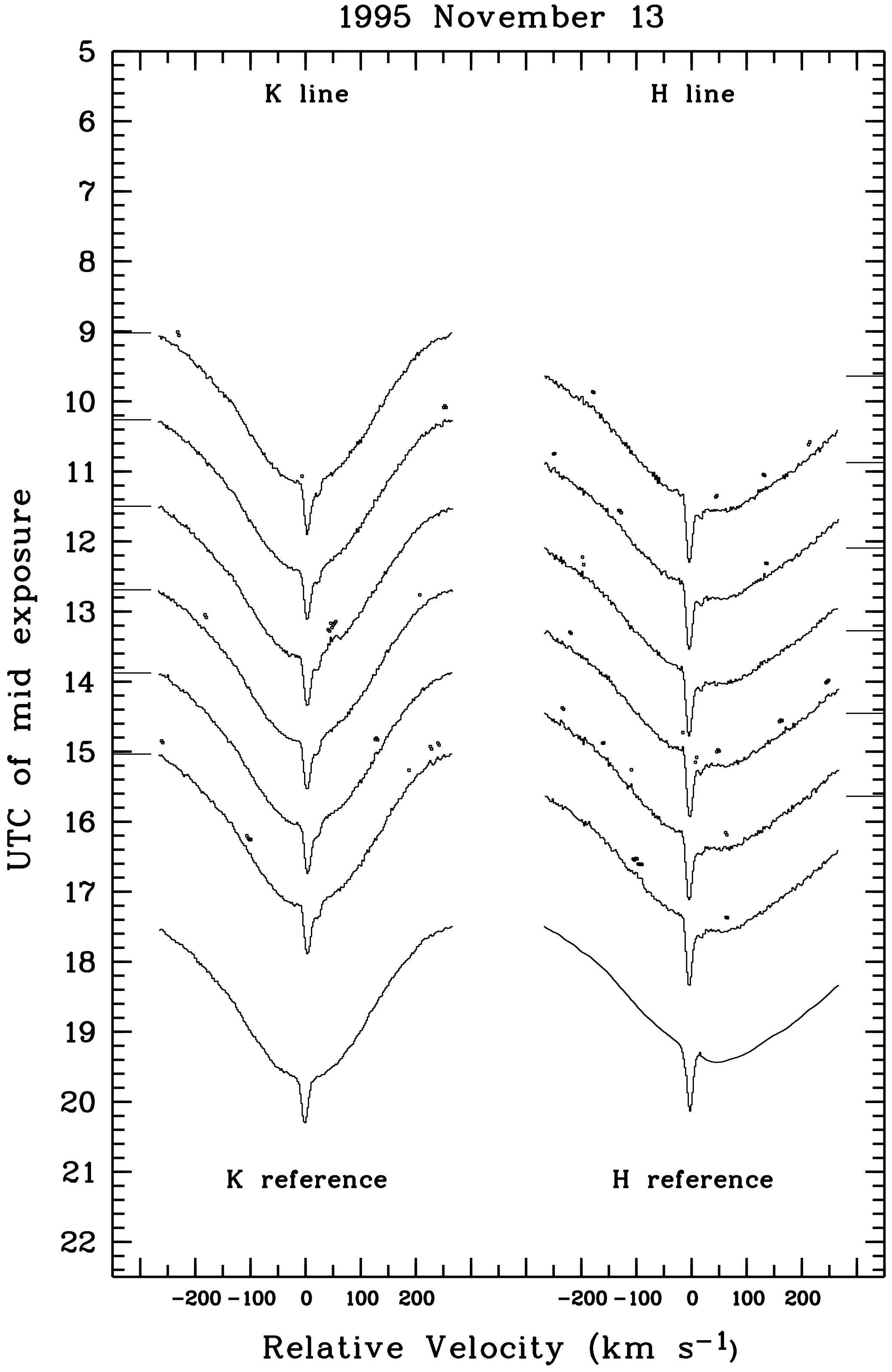

Figure 37: Observations of β Pictoris on 1995 November 13 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

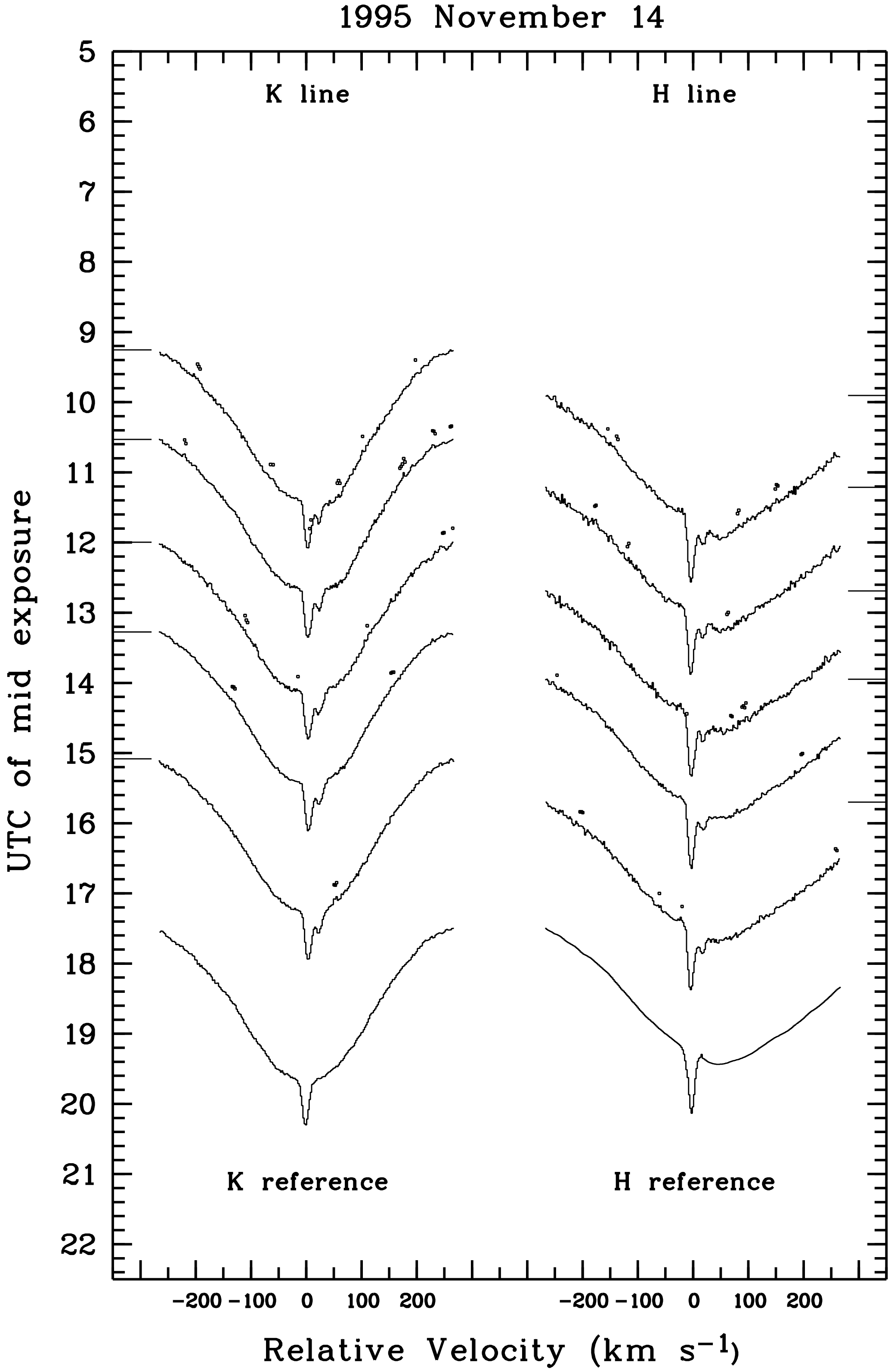

Figure 38: Observations of β Pictoris on 1995 November 14 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

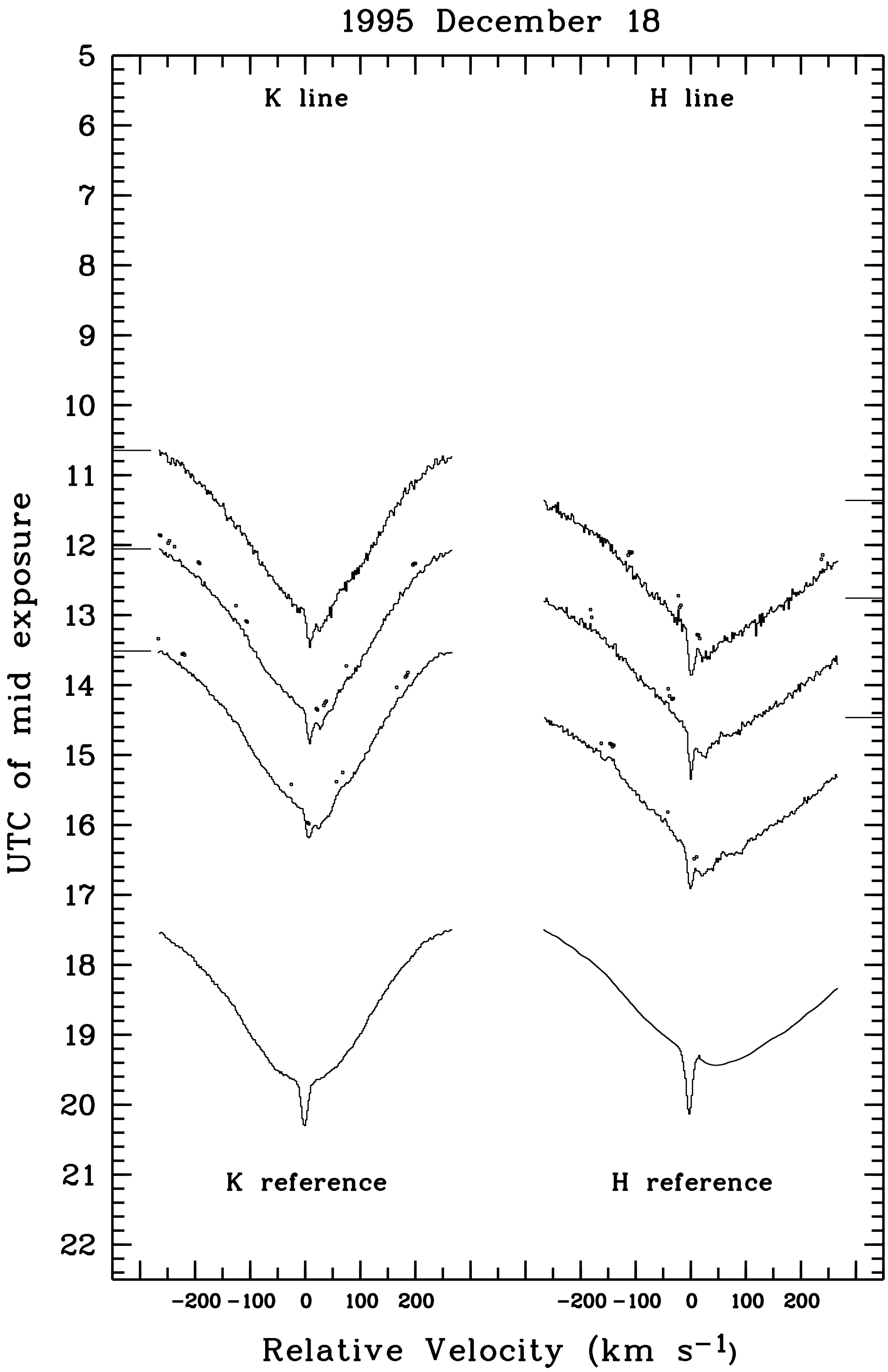

Figure 39: Observations of β Pictoris on 1995 December 18 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.

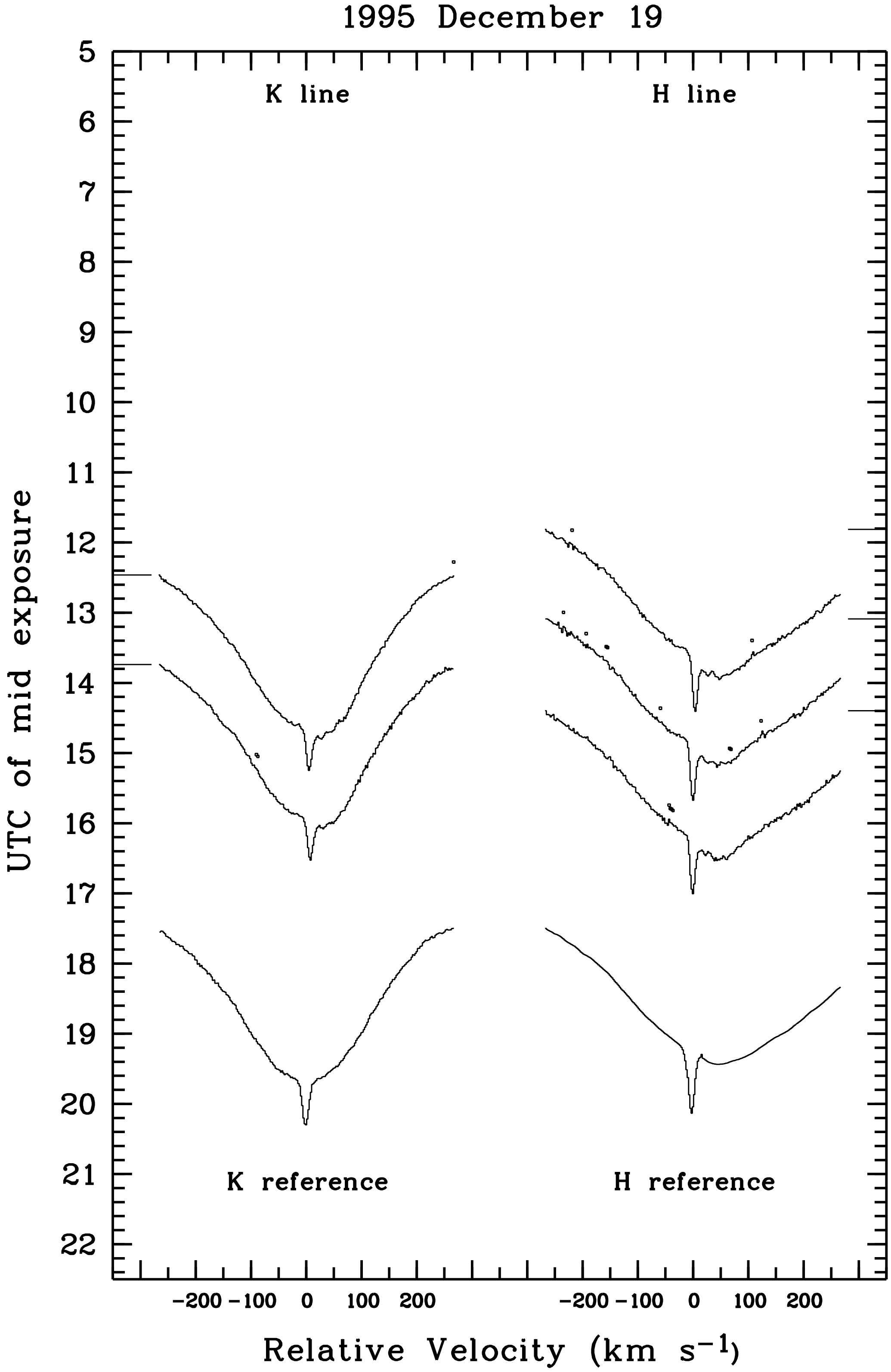

Figure 40: Observations of β Pictoris on 1995 December 19 UTC. Dots above spectra indicate where cosmic rays have been detected and filtered out. The marks on the ordinate scale indicate the time of mid exposure. The velocity zero points are only approximately centred on the circumstellar features.