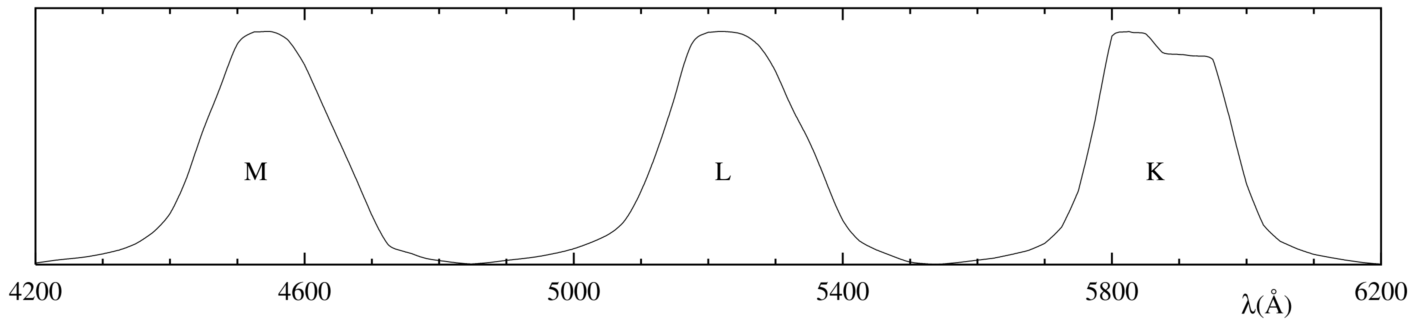
KLMNPQR - Borgman - 1960 (continues)



M				*B3*	*Vega*	*Sun*	*K2*	*M2*	*Carbon*
$\lambda_c = 4548$	$\lambda_\circ = 4542$	$\lambda_{peak} = 4545$	$\lambda_{gauss} = 4546$	4537	4539	4546	4553	4558	4529
WHM = 216	W10% = 365	W80% = 130	FWHM = 215	[227]	[226]	[226]	[223]	[219]	[196]
$W_\circ = 229$	$\frac{A(\lambda)}{A(V)}\vert_{5.0} = 1.18\ {}^{1.17}_{1.22}$	$a = {}^{1.006}_{1.006}$ $b = {}^{0.833}_{0.823}$	*B3*	*WN*	*WC*	PN_{Ne}	PN^{Ne}	*Nova*	*WDA*
$\mu = 91$	$\frac{A(\lambda)}{A(V)}\vert_{3.1} = 1.29\ {}^{1.27}_{1.36}$	$a = {}^{1.006}_{1.007}$ $b = {}^{0.821}_{0.812}$	*Sun*	4570	4595	4537	4537	4552	4540
$I_{asym} = -0.27$	$\frac{A(\lambda)}{A(V)}\vert_{2.1} = 1.43\ {}^{1.40}_{1.54}$	$\frac{A(\lambda)}{E(B-V)}$:	$(3.947, 0.070)^{r=1.00}_{B3}$	$(4.231, 0.089)^{r=1.00}_{Sun}$			$(4.819, 0.075)^{r=1.00}_{M2}$		
$I_{kurt} = 0.50$	$\lambda_{eff} = 4542.3 + 7.8 \times E(B-V)$ r=1.00			$W_{eff} = 228.4 - 2.4 \times E(B-V)$ r=−0.96					
	$\lambda_{eff}(T) = 4535 + 16 \times \theta + 11 \times \theta^2 - 6 \times \theta^3$			$W_{eff}(T) = 226 + 6 \times \theta - 8 \times \theta^2 - 1 \times \theta^3$					

L				*B3*	*Vega*	*Sun*	*K2*	*M2*	*Carbon*
$\lambda_c = 5240$	$\lambda_\circ = 5231$	$\lambda_{peak} = 5219$	$\lambda_{gauss} = 5235$	5225	5226	5231	5239	5251	5271
WHM = 229	W10% = 395	W80% = 147	FWHM = 230	[243]	[243]	[245]	[245]	[234]	[178]
$W_\circ = 245$	$\frac{A(\lambda)}{A(V)}\vert_{5.0} = 1.05\ {}^{1.04}_{1.08}$	$a = {}^{1.012}_{1.012}$ $b = {}^{0.153}_{0.148}$	*B3*	*WN*	*WC*	PN_{Ne}	PN^{Ne}	*Nova*	*WDA*
$\mu = 96$	$\frac{A(\lambda)}{A(V)}\vert_{3.1} = 1.08\ {}^{1.06}_{1.12}$	$a = {}^{1.012}_{1.012}$ $b = {}^{0.149}_{0.144}$	*Sun*	5230	5230	5051	5225	5117	5226
$I_{asym} = -0.31$	$\frac{A(\lambda)}{A(V)}\vert_{2.1} = 1.11\ {}^{1.09}_{1.17}$	$\frac{A(\lambda)}{E(B-V)}$:	$(3.288, 0.058)^{r=1.00}_{B3}$	$(3.526, 0.074)^{r=1.00}_{Sun}$			$(4.006, 0.061)^{r=1.00}_{M2}$		
$I_{kurt} = 0.46$	$\lambda_{eff} = 5231.2 + 7.8 \times E(B-V)$ r=1.00			$W_{eff} = 244.7 - 2.5 \times E(B-V)$ r=−0.97					
	$\lambda_{eff}(T) = 5224 + 13 \times \theta + 10 \times \theta^2 - 5 \times \theta^3$			$W_{eff}(T) = 243 + 4 \times \theta - 4 \times \theta^2 - 2 \times \theta^3$					

K				*B3*	*Vega*	*Sun*	*K2*	*M2*	*Carbon*
$\lambda_c = 5876$	$\lambda_\circ = 5874$	$\lambda_{peak} = 5825$	$\lambda_{gauss} = 5872$	5870	5870	5873	5875	5871	5878
WHM = 220	W10% = 347	W80% = 172	FWHM = 219	[232]	[232]	[232]	[232]	[236]	[231]
$W_\circ = 233$	$\frac{A(\lambda)}{A(V)}\vert_{5.0} = 0.94\ {}^{0.94}_{0.97}$	$a = {}^{0.973}_{0.972}$ $b = {}^{-0.136}_{-0.138}$	*B3*	*WN*	*WC*	PN_{Ne}	PN^{Ne}	*Nova*	*WDA*
$\mu = 87$	$\frac{A(\lambda)}{A(V)}\vert_{3.1} = 0.91\ {}^{0.90}_{0.96}$	$a = {}^{0.972}_{0.972}$ $b = {}^{-0.137}_{-0.138}$	*Sun*	5869	5832	5865	5872	5848	5870
$I_{asym} = 0.04$	$\frac{A(\lambda)}{A(V)}\vert_{2.1} = 0.87\ {}^{0.85}_{0.93}$	$\frac{A(\lambda)}{E(B-V)}$:	$(2.784, 0.050)^{r=1.00}_{B3}$	$(2.989, 0.064)^{r=1.00}_{Sun}$			$(3.416, 0.054)^{r=1.00}_{M2}$		
$I_{kurt} = 0.31$	$\lambda_{eff} = 5874.3 + 5.1 \times E(B-V)$ r=1.00			$W_{eff} = 232.3 - 1.3 \times E(B-V)$ r=−0.97					
	$\lambda_{eff}(T) = 5869 + 8 \times \theta + 7 \times \theta^2 - 3 \times \theta^3$			$W_{eff}(T) = 231 + 2 \times \theta - 2 \times \theta^2 - 1 \times \theta^3$					